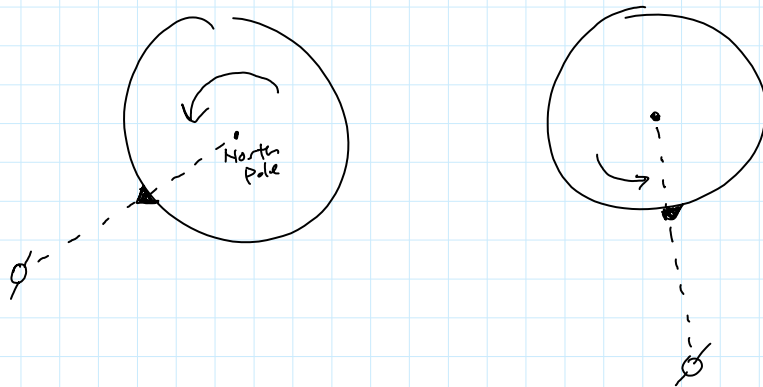


Goals for the Lecture:

- 1) Understand what Period, Frequency, and angular frequency are and how they relate to one another
- 2) Understand how velocity and acceleration change during a cycle
- 3) Understand that simple harmonic motion comes about when the force is proportional to displacement

Geosynchronous orbit



Find R for geosynchronous orbit:

$$\sum \vec{F}_{\text{radial}} = m a_c$$

$$\sum \vec{F}_{\text{radial}} = m \frac{v^2}{R}$$

$$\frac{G M_E m}{R^2} = m \frac{v^2}{R}$$

$$v = \frac{\text{distance}}{\text{time}} = \frac{2\pi R}{T}$$

↑
period = 24 hours

$$\frac{G M_E}{R^2} = \frac{(2\pi R)^2}{T^2 R}$$

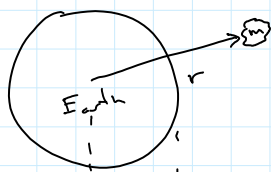
$$\frac{G M_E}{4\pi^2} T^2 = R^3$$

↑
 $T = 1 \text{ day}$

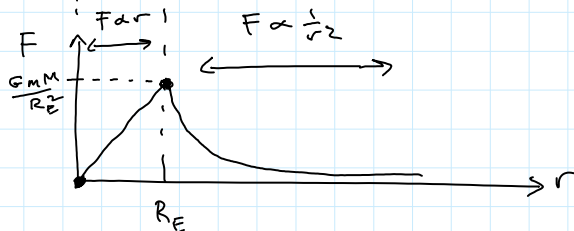
solve for R

Gravity:

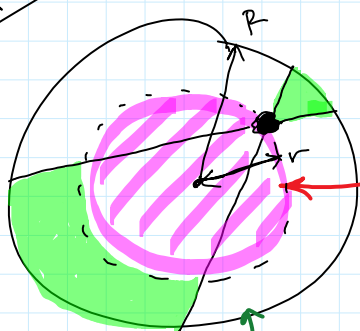
For $r > R_E$



$$F = \frac{G m M_E}{r^2}$$



For $r < R$



only mass inside our radius has an effect

all mass outside our radius cancels out

$$F = \frac{G m M_{\text{inside}}}{r^2}$$

if uniform density:

$$M_{\text{inside}} = \frac{V_{\text{inside}}}{V_{\text{total}}} M_{\text{total}}$$

$$= \frac{\frac{4}{3}\pi r^3}{\frac{4}{3}\pi R^3} M$$

$$= \frac{r^3}{R^3} M$$

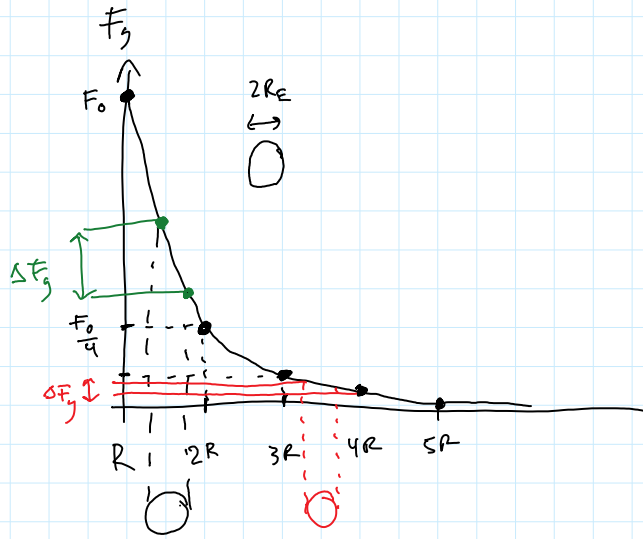
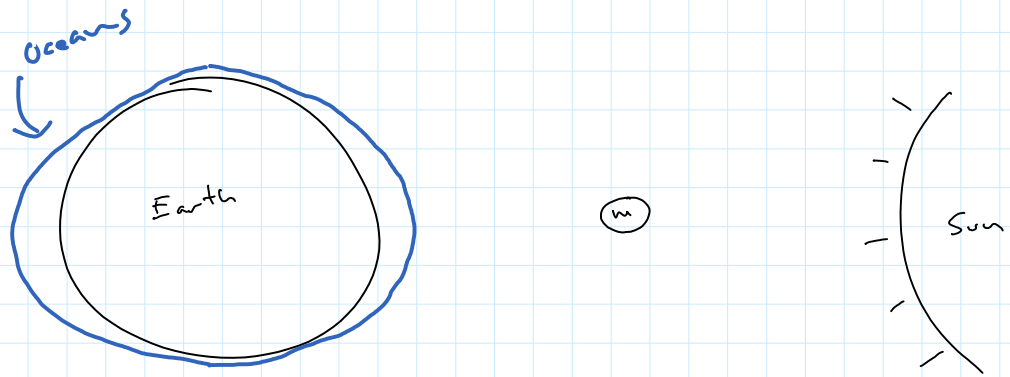
$$F = \frac{G m \left(\frac{r^3}{R^3} M_{\text{total}} \right)}{r^2}$$

$$= \left(\frac{G m M_{\text{total}}}{R^3} \right) r$$

$F \propto r$ inside planet

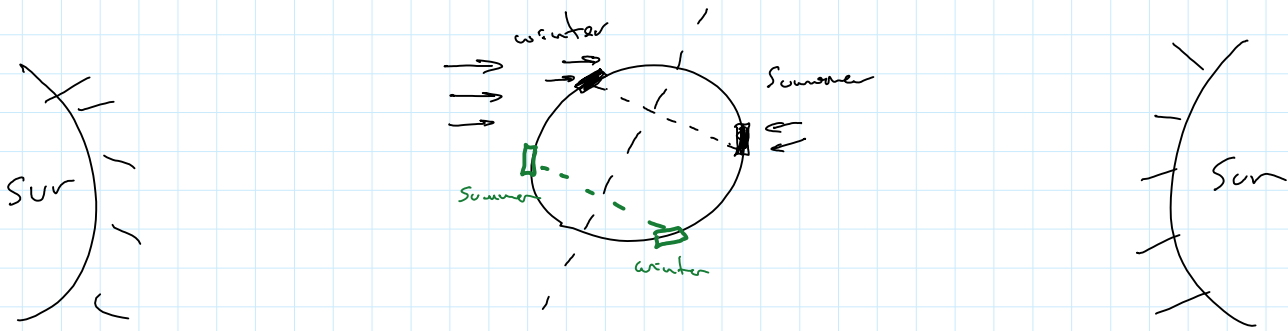
$F \propto r$ inside planet

Tides



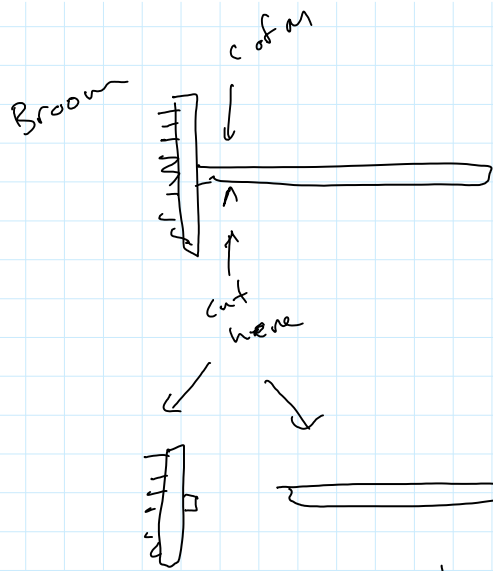
$$F = \frac{GmM}{r^2}$$

Seasons:



Flux

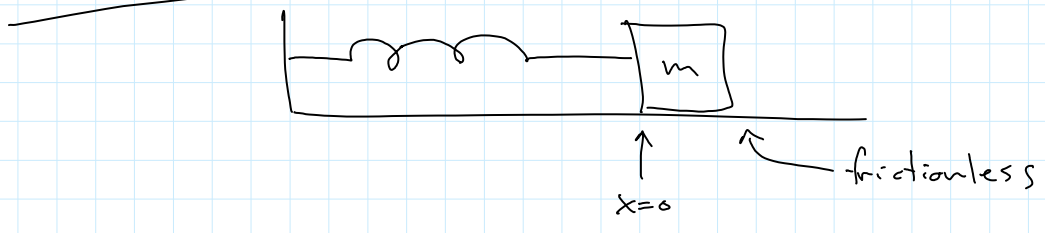
c of m



which is heavier?
 Broom end is heavier because it has a smaller lever arm

$$\sum_{\text{right}} = \sum_{\text{left}}$$

Mass on a Spring:



$$F = -kx$$

$$\sum F = ma$$

$$-kx = ma$$

$$-kx = m \frac{d^2x}{dt^2}$$

$$-\frac{k}{m} x = \frac{d^2x}{dt^2}$$

$$x(t) = ?$$

if $x(t) = A \cos(\omega t + \phi)$

for setting initial conditions

$$\text{if } x(t) = A \cos(\omega t + \phi)$$

$$x'(t) = -\omega A \sin(\omega t + \phi) = v(t)$$

$$x''(t) = \underbrace{-\omega^2 A}_{-\frac{F}{m}} \underbrace{\cos(\omega t + \phi)}_{x(t)} = a(t)$$

$$\text{if } \omega = \sqrt{\frac{k}{m}}$$

$$x(t) = A \cos(\omega t + \phi) \text{ satisfies}$$

ω = angular frequency units: $\frac{\text{radians}}{\text{sec}}$

$$\omega = 2\pi f$$

↑

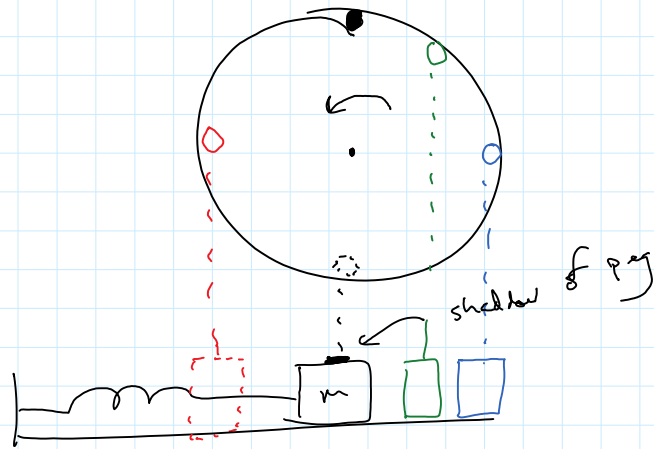
frequency

f units

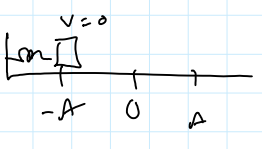
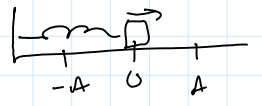
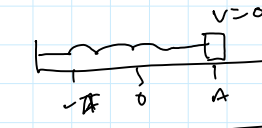
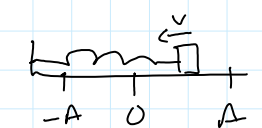
$\frac{\text{cycles}}{\text{sec}}$ or Hertz (Hz)

Light ↓ ↓ ↓ ↓ ↓ ↓

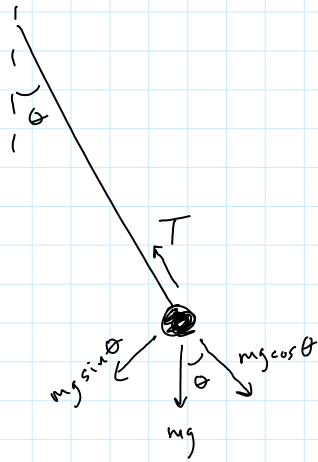
$$\text{Hz} = \frac{1}{\text{sec}}$$



	t	x	v	a	KE	U _{sp}	
	0	A	0	$-\frac{k}{m}A$ OR $-\omega^2 A$	0	$\frac{1}{2}kA^2$	$\phi = 0$
	$\frac{T}{4}$	0	$-\omega A$	0	$\frac{1}{2}kA^2$	0	
	T	-A	0	$+\frac{k}{m}A$	0	$\frac{1}{2}kA^2$	

	$\frac{T}{2}$	$-A$	0	$+\frac{kA}{2}$ or $\omega^2 A$	0	$\frac{1}{2}kA^2$
	$\frac{3}{4}T$	0	ωA	0	$\frac{1}{2}kA^2$	0
	T	A	0	$-\omega^2 A$	0	$\frac{1}{2}kA^2$
	$\frac{5}{4}T$	x	$-v$	$-\omega^2 x$	$\frac{1}{2}mv^2$	$\frac{1}{2}kx^2$

simple pendulum



$$F = ma$$

$$-mg \sin \theta = ma$$

$$-g \sin \theta = a$$

if θ is small: $\sin \theta \approx \theta$ in radians

$$-g \theta = a$$

$$a = \frac{d^2 s}{dt^2}$$

$$s = r\theta$$

$$-\frac{g}{l} \theta = \frac{d^2 \theta}{dt^2}$$

$$\theta(t) = A \cos(\omega t + \phi)$$

$$\omega = \sqrt{\frac{g}{L}}$$